An On Line Interactive Geometric Database: Including Exact Solutions of Einstein's Field Equations

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We describe a new interactive database (GRDB) of geometric objects in the general area of differential geometry. Database objects include, but are not restricted to, exact solutions of Einstein's field equations. GRDB is designed for researchers (and teachers) in applied mathematics, physics and related fields. The flexible search environment allows the database to be useful over a wide spectrum of interests, for example, from practical considerations of neutron star models in astrophysics to abstract space-time classification schemes. The database is built using a modular and object-oriented design and uses several Java technologies (e.g. Applets, Servlets, JDBC). These are platform-independent and well adapted for applications developed to run over the World Wide Web. GRDB is accompanied by a virtual calculator (GRTensorJ), a graphical user interface to the computer algebra system GRTensorII used to perform on line coordinate, tetrad or basis calculations. The highly interactive nature of GRDB allows for systematic internal self-checking and a minimization of the required internal records. This new database is now on line at http://grdb.org.

I. INTRODUCTION

The study of exact solutions of Einstein's Field Equations is an important part of the theory of General Relativity [1] [2]. This importance derives not only from more formal mathematical aspects associated with the theory (e.g. the classification of space-times) but also from the growing importance of the application of general relativity to astrophysical phenomena. For example, exact solutions can offer physical insights that numerical solutions can not. There are of course a very large number of exact solutions offered in the literature. However, these include numerous re-discoveries, confusions and errors. In a recent study of static spherically symmetric perfect fluid solutions, for example, only 9 of 127 candidate solutions were found to be of physical interest [3].

It is not surprising that some of the earliest applications of computer algebra can be found in General Relativity [4] [5]. Indeed, no effective study of exact solutions in the large can be carried on without it. With the recent rapid development of the World Wide Web (WWW) as an means of information transfer, it has become an effective medium for cataloguing exact solutions. The first such use of the WWW is due to J. Skea [6]. Skea's database is primarily a classification database [7]. It is static in the sense that the information that can be retrieved is resident in the records.

In this paper we report on a non-static WWW based database, one for which the information that can be retrieved is not restricted to information resident in the records. This development has become feasible using more recent WWW technologies. This new database is now on line at http://grdb.org (henceforth GRDB) using several Java technologies and following a fully modular object-oriented design. The dynamic nature of the database is accomplished via the inclusion of GRTensorJ [8]. This is an interactive programmable graphical user interface to the computer algebra system GRTensorII [9] which in turn runs under Maple [10].

GRDB is designed to be a database of geometric objects in the general area of differential geometry with no a priori restriction on the number of dimensions. It contains exact solutions of Einstein's field equations without being restricted to them. By design the database is highly interactive. The records (manifolds) are generally stored only in terms of a metric or a set of basis vectors along with constraints and referencing data. Other elements can be calculated and displayed as required either interactively or from an extensive menu.

The database is designed for researchers (and teachers) in applied mathematics, physics and related fields and it is hoped that the database will find use over a wide spectrum of interests.

II. DESCRIPTION AND USAGE

GRDB offers many choices of operation. These operations are typically carried in the following order:

• i) Search the database by fields of particular interest. A search by keywords or by an equation number e.g. as listed in the "Exact Solutions Book" [1] is implemented. Other searches can be easily added as the project evolves, e.g. a search by the equation number as listed in any standard reference. The keyword choice searches the information field and the name of the metric field of each entry in the database. A common use is a search based on an author's surname (e.g. Kruskal). The key word search must be at least 3 characters. Reference searches follow a simple referencing protocol. The search result is displayed in a list as shown on the bottom

- of Figure 1. The record short name is displayed as a hyper-link. The record naming convention is explained in Appendix A.
- ii) Select a record or records of interest from the search result list by clicking on its hyper-link. This will display a detailed description of the space-time starting with a suggestive line element representation of the record that is produced dynamically (not stored). Record details are then displayed as in Figure 2 with a structure adapted to the type of the record examined (i.e. a metric, a tetrad or a basis type entry).
- iii) Mark the viewed record(s) to be loaded into the on line calculator GRTensorJ. This calculator is described further bellow.
- iv) Click on the Start Calculator button from the database home page in order to launch GRTensorJ and to do on line calculations for the loaded records.
- v) Click on "Load/Select" from the menu inside the calculator (Figure 3) and then click on "Load Marked Manifold(s) from Database". The tagged manifolds will appear as choices under Load/Select. Click on the manifold you wish to use for calculations. This manifold is now loaded into the calculator and the user can now perform calculations. An extensive list of objects has already been implemented. This is summarized in Appendix B. Customized calculations are described below.
- vi) Select the calculations commands through menu and sub-menu selections in the calculator as shown in Figure 3. The first step is to select a space-time in coordinates, tetrads or basis. This choice automatically displays the components of the metric tensor, the tetrads, or the basis components. Then, the user can select the object(s) to be calculated from the corresponding menus. After the result of the calculation is displayed the user can apply to it any simplification procedure supported by the engine as listed under "Simplifications" in the menu bar. A help system is built into the menu system. Further, the user can define new objects through the definition facilities within GRTensorII. These are entered through a sub-interface and saved for the user's work session. There are also other support and option selection functions. It is especially important to note that if a manifold is represented with constraints (e.g Kruskal Szekeres) then the constraints (under "Simplifications") must be applied on the calculations, otherwise the user is not working with the selected manifold. A number of useful features are easily added to GRTensorJ. For example, automatic Latex has been implemented [11].
- vii) New objects are defined in a sub-menu using the syntax of the GRTensorII grdef facility [9]. For example, the Bel-Robinson tensor (though pre-defined) would be entered as $T\{(cdef)\} := C\{acdb\} * C\{aefb\} + Cstar\{acdb\} * Cstar\{aefb\} *$. In addition, there are objects that require an input from the user, e.g. the kinematical quantities associated with a field u^a . In order to create an object, select a manifold, then select under the menu item "Custom" the option "Create a new single definition object". Follow the syntax explained under "Explanation" and the example shown on the object input interface, e.g. ` $u^a := [u^1, u^2, u^3, u^4]$ `where u^1 etc. are the components. Once the object is created, it is saved during the user's work session and related calculations can be performed, e.g. the acceleration, shear and rotation for u^a . For a full description of the syntax the user may consult GRTensorII [9] documentation on the command "grdef" or browse the on line help system from the GRTensorII WWW site [9].
- viii) Input in the database. There are two forms of input into the database, a Public Input and a Private Input. It is our intention that all public input be refereed. Private input is password protected and reserved for maintenance and internal operations. The List-All option is currently for private usage. The public input includes three types of interfaces; a line element, a covariant or contravariant Newman-Penrose tetrad or a non-holonomic basis. Figure 4 gives an example of a line element input interface.

III. INTERNAL DESIGN

A. GRDB components

GRDB is built using fully modular Client/Server and object-oriented designs. It uses Java Servlet technology [12], a platform-independent mechanism for programming applications over the WWW. Java is a pure object-oriented language, allowing multi-thread processes, specially adapted to network applications. Java Servlets are secure and portable applications that handle requests from a Web server that are originating from a client Web browser. The servlet runs on the server side and does not depend on the browser compatibility. Java Servlets also solve many access issue problems, as they can be used to develop fully distributed applications over the Web where the database

can be on separate machines. GRDB also uses a JDBC (Java DataBase Connectivity) module [13], a portable and flexible application that acts as an independent connector to the database, increasing the modularity of the architecture. PostgreSQL [14] was chosen and configured as the DataBase Management System (DBMS), a system that is also object-oriented and that brings to GRDB all the power and performance of the Structured Query Language (SQL) commands. This language makes any kind of search or data manipulation very easy to implement. The SQL instructions are formulated within the Java Servlet code. The web-pages used in order to manipulate the Data are dynamic pages created while the programs are running. Only the help pages were left static. This HTML is also embedded in the Servlet code. The records (manifolds) are for the most part stored only in terms of a metric or set of basis vectors along with constraints and referencing data. This information is processed interactively through the Servlets and GRTensorJ which in turn runs on Maple. The program TTH [15] is used to translate TEX into HTML in order to produce mathematical html text.

B. The Calculator GRTensorJ

GRTensorJ is also written in Java and has been designed using a full object-oriented approach. In addition, all communication between the user and the server is object-based. GRTensorJ has a multi-layer architecture that allows the generic functionality described above. This architecture is outlined below:

- GUI (Graphical User Interface)
- User Functional Interface
- User ICM (Interchange Module) Handler
- Interchange Modules
- Server ICM Handler
- Server Functional Interface
- Algebraic Engine Server Structure

A Java Applet (a programmed application that runs on the client side) opens GRTensorJ by clicking on the applet button on GRDB home page. It appears as in Figure 3. Behind the scenes, a computer algebra session in GRTensorII is started automatically. GRTensorJ uses Maple [10] as the algebraic engine. GRTensorJ is designed with an open architecture and is compatible with any other algebraic engine that can output an ASCII stream. This architecture also allows it to be expanded and programmed by the user. The commands are accessed through menu and submenu selections. When a session begins, GRTensorJ reads a directory on the server named TextSheets (not to be confused with "worksheets") and builds the menu and sub-menus for the interface from the underlying structure. All sub-directories to TextSheets will appear as primary menu bar items. The names of ASCII files contained within these sub-directories will be displayed as menu selections. These files contain a sequence of commands written in the syntax of the computer algebra engine being used. By selecting a menu item the user sends these commands to the engine. Items to be displayed in the interface window are distinguished simply by an asterisk in the file. In other words, creating new menu items and calculation commands is as simple as creating and editing very simple ASCII files. Yet, the file, and resultant menu item, can be the equivalent of an entire worksheet with only the result chosen for display. Of course, users need program nothing at all. The underlying simplicity of GRTensorJ is mentioned here simply to demonstrate its flexibility.

IV. FURTHER DEVELOPMENTS

When first released (July 2001), the database contained only about 300 entries. This is under continuous development and, for example, it is planed to have the entire content of [1] and [2] included in the database. An extension of the database to incorporate algorithmic classification is in progress [16] as is the construction of specialized record data associated with, for example, static spherically symmetric perfect fluid solutions [3].

Acknowledgments

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- [9] This is a package which runs within Maple. It is entirely distinct from packages distributed with Maple and must be obtained independently. The GRTensorII software and documentation are distributed freely on the World-Wide-Web from http://grtensor.org and http://www.aei-potsdam.mpg.de/~pollney/grtensor/
- [10] Maple is copyright Waterloo Maple Software and use of Maple through GRTensorJ on the Web is by special arrangement with Waterloo Maple Software.
- [11] A latex call in Maple produces output suitable for use with LaTeX 2e and appears in the applet window. We are serious when we say that you can, for example, now go to the Web, do your calculation, and paste the answer into the paper you are writing!
- [12] See http://java.sun.com/products/servlets.
- $[13] \operatorname{See} \text{ http://java.sun.com/products/jdbc.}$
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APPENDIX A: RECORD NAMING CONVENTION

The record naming convention is not of significant interest to the user as it is merely a placeholder for links. The convention does, however, convey some information that can be quickly digested by the user. The naming convention is as follows: "abc"

- 1: No numbers
- 2: All lower case
- 3: a=name (e.g. bondi)
- 4: b=m (metric), or npd (covariant tetrad), or npu (contravariant tetrad), or b (basis)
- 5: c=a,...,z (version)
- 6: Default signature for metrics +2 for NP tetrads -2.

examples:

kerrma means the Kerr spacetime in metric form version a kerrmb means the Kerr spacetime in metric form version b kerrnpda means the Kerr spacetime in NP tetrad covariant form version a kerrnpua means the Kerr spacetime in NP tetrad contravariant form version a.

APPENDIX B: INITIAL SET OF PREPROGRAMMED OBJECTS AVAILABLE FOR CALCULATION

GRTensorJ menu item GRTensorJ sub-menu command Object(s)

Metric	Metric	g_{ab}
	Signature	Sig
	Line Element	ds^2
	Determinant of Metric	det(g)
	Inverse Metric	g^{ab}
	Partial Derivative of Metric	$g_{ab,c}$
Christoffel Symbols	Chr(dn,dn,dn)	Γ_{abc}
	Chr(dn,dn,up)	Γ^c_{ab}
Geodesic		Writes out the geodesic equations
Riemann	R(dn,dn,dn,dn)	R_{abcd}
	R(up,dn,dn,dn)	$R^a_{\ bcd}$
	R(up,up,dn,dn)	R^{ab}_{cd}
	R(up,up,up,up)	R^{abcd}
	Kretschmann	$R_{abcd}R^{abcd}$
Weyl	C(dn,dn,dn,dn)	C_{abcd}
-	Dual CStar(dn,dn,dn,dn)	$C^*_{abcd} \equiv \frac{1}{2} \epsilon_{abef} C^{ef}_{cd}$
Ricci	R(dn,dn)	$R_{ab} \equiv R^{c}_{acb}$
Trace-free Ricci	S(dn,dn)	$S_{ab} \equiv R_{ab} - \frac{1}{N} q_{ab} R$
Einstein	G(dn,dn)	$G_{ab} \equiv R_{ab} - \frac{1}{2}q_{ab}R$
Invariants	Invariants-Ricci	$S_{ab} \equiv R_{ab} - \frac{1}{N} g_{ab} R$ $G_{ab} \equiv R_{ab} - \frac{1}{2} g_{ab} R$ $R \equiv R^a_{\ a}, R1 \equiv \frac{1}{4} S^a_{\ b} S^b_{\ a}, R2 \equiv \frac{-1}{8} S^a_{\ b} S^b_{\ c} S^c_{\ a}$
		$R3 \equiv \frac{1}{12} S^a {}_b S^b {}_c S^c {}_d S^d {}_a$
	Invariants-Weyl	$W1R = \frac{1}{2}C_{-1-1}C^{abcd}$ $W1I = \frac{1}{2}C^*$
	THE CONTRACTOR OF THE CONTRACT	$W2R = \frac{-1}{2}C$, cdC , efC , ab
		$W2I = \frac{1}{16}Cab$ Ca Ce I Ce I
	Invariants-Mixed	$R \equiv R^{-}_{a}, K1 \equiv \frac{1}{4}S^{-}_{b}S^{-}_{a}, K2 \equiv \frac{1}{8}S^{-}_{b}S^{-}_{c}S^{-}_{a}$ $R3 \equiv \frac{1}{16}S^{a}_{b}S^{b}_{c}S^{c}_{d}S^{d}_{a}$ $W1R \equiv \frac{1}{8}C_{abcd}C^{abcd}, W1I \equiv \frac{1}{8}C^{*}_{abcd}C^{abcd}$ $W2R \equiv \frac{-1}{16}C_{ab}^{cd}C_{cd}^{ef}C_{ef}^{ab}$ $W2I \equiv \frac{-1}{16}C^{*}_{ab}^{cd}C_{cd}^{ef}C_{ef}^{ab}$ $M1R \equiv \frac{1}{8}S^{ab}S^{cd}C_{abcd}, M1I \equiv \frac{1}{8}S^{ab}S^{cd}C^{*}_{abcd}$ $M2R \equiv \frac{1}{16}S^{cd}S_{ef}(C_{acdb}C^{aefb} - C^{*}_{acdb}C^{aefb})$ $M2I = \frac{1}{16}S^{cc}S_{ef}(C_{acdb}C^{aefb} - C^{*}_{acdb}C^{aefb})$
	THVallands Whaca	$M1R = \frac{1}{8}S^{cd}S$ (C :: $C^{aefb} = C^*$ C^{aefb})
		$M2It = \frac{1}{16}S S_{ef}(C_{acdb}C - C_{acdb}C_*)$ $M2I = \frac{1}{8}S^{bc}S_{ef}(C_{abcd}^*C^{aefd})$
		$M2I \equiv \frac{1}{8}S Sef(C_{abcd}C)$ $M2 = \frac{1}{8}S Gef(C_{abcd}C) G^{aefb} G^* G^{aefb}$
		$M3 = \frac{1}{16}S$ $S_{ef}(C_{acdb}C + C_{acdb}C_*)$
		$M3 \equiv \frac{1}{16} S^{cd} S_{ef} (C_{acdb} C^{aefb} + C^*_{acdb} C^{aefb})$ $M4 \equiv \frac{-1}{32} S^{cg} S^{ef} S^{c}_{d} (C_{ac}^{db} C_{befg} + C^*_{ac}^{db} C^*_{befg})$ $M5R \equiv \frac{1}{32} S^{cd} S^{ef} C^{aghb} (C_{acdb} C_{gefh} + C^*_{acdb} C^*_{gefh})$ $M5I \equiv \frac{1}{32} S^{cd} S^{ef} C_*^{aghb} (C_{acdb} C_{gefh} + C^*_{acdb} C^*_{gefh})$ $M6I \equiv \frac{1}{32} S^{cd} S^{ef} C_*^{aghb} (C_{acdb} C_{gefh} + C^*_{acdb} C^*_{gefh})$
		$M5R \equiv \frac{1}{32}S^{as}S^{as}C^{ash}(C_{acdb}C_{gefh} + C_{acdb}C_{gefh})$
		$M5I \equiv \frac{1}{32}S^{cd}S^{ef}C_*^{aghb}(C_{acdb}C_{gefh} + C_{acdb}C_{gefh})$
		$M6R \equiv \frac{1}{32} S_a{}^e S_e{}^c S_b{}^J S_f{}^d C^{ab}{}^{cd}$
		$M6I \equiv \frac{1}{32} S_a^{\ e} S_e^{\ c} S_b^{\ j} S_f^{\ d} C^{*ab}_{\ cd}$
Differential Invariants	diRicci	$MSI = \frac{3}{32}S S^{c} S^{c} C_{*} C_{acdb} C_{gefh} + C_{acdb} C_{gefh})$ $M6R = \frac{1}{32}S_{a}^{e} S_{e}^{c} S_{b}^{f} S_{f}^{d} C^{ab}_{cd}$ $M6I = \frac{1}{32}S_{a}^{e} S_{e}^{c} S_{b}^{f} S_{f}^{d} C^{*ab}_{cd}$ $R_{ab;c} R^{ab;c}$
	diRiem	$R_{abcd:e}R^{abcd;e}$
	diS	$S_{ab;c}S^{ab;c}$
	diWeyl	$C_{abcd;e}C^{abcd;e}$
Bel-Robinson	$T(dn,dn,dn,dn) \dots$	$T_{cdef} \equiv C_{acdb}C^{a}_{ef}^{b} + C^{*}_{acdb}C^{*a}_{ef}^{b}$
Weyl-Schouten	Weyl-Schouten	
Bach	B(dn,dn)	$WS_{abc} \equiv R_{ab;c} - R_{ac;b} - (g_{ab}R^e_{e;c} - g_{ac}R^f_{f;b})/4$ $B_{ac} \equiv C_{abcd}^{;bd} + \frac{1}{2}R^{bd}C_{abcd}$
Input field calculations	Field u	u^a
•	KillingCoords	Adapted coordinate check
	KillingTest	Test for conformal/homothetic/Killing vector
	Vector Kinematics	$\dot{u}^a \equiv u^b u^a_{;b}, \ \theta \equiv u^b_{;b}, \ \sigma_a{}^b \equiv u_{(a}{}^{;b}{}_) + \dot{u}_{(a} u^b{}_) -$
	Electric Weyl E(dn,dn)	$ \frac{1}{3}h_a{}^b\dot{\theta}, \omega_{ab} \equiv u_{[a;b]} + \dot{u}_{[a}u_{b]} $ $ E_{ac} \equiv C_{abcd}u^b u^d $
	Magnetic Weyl H(dn,dn)	$H_{ac} \equiv C^*_{abcd} u^b u^d$
		and — C aucaw w

TABLE I: Initial set of preprogrammed objects available for coordinate calculations. Note: "..." means that all index combinations are immediately available as with Riemann.

GRTensorJ menu item	GRTensorJ sub-menu command	Object(s)
NPSpin	NPSpin: NPkappa, NPsigma, NPlambda, NPnu, NPrho, NPmu, NPtau, NPpi, NPepsilon, NPgamma, NPalpha, NPbeta	spin coefficients, $\kappa, \sigma, \dots, \beta$
NPSpinbar	NPSpinbar: NPkappabar, NPsigmabar, NPlambdabar, NPnubar, NPrhobar, NPmubar, NPtaubar, NPpibar, NPepsilonbar, NPgammabar, NPalphabar, NPbetabar	spin coefficients (complex conjugates) $\bar{\kappa}, \bar{\sigma}, \dots \bar{\beta}$
Ricci Scalars	Ricci Scalars: Phi00, Phi01, Phi02, Phi10, Phi11, Phi12, Phi20, Phi21, Phi22, Lambda	Ricci scalars, $\Phi_{00}, \Phi_{01}, \dots, \Phi_{22}, \Lambda$
Weyl Scalars	Weyl Scalars: Psi0, Psi1, Psi2, Psi3, Psi4	Weyl scalars, Ψ_0, \dots, Ψ_4
	Petrov Petrov Report	Petrov type A report on how the Petrov type was determined

TABLE II: Initial set of preprogrammed objects available for tetrad calculations. All the objects in TABLE I are integrated with tetrad input and can be run directly from a tetrad.

GRTensorJ menu item GRTensorJ sub-menu command Object(s)

Rotation Coefficients	lambda(bdn,bdn,bdn) rot(bdn,bdn,bdn)	$\lambda_{(a)(b)(c)} := e_{(b)[i,j]} e_{(a)}^{d} e_{(c)}^{e}$ rotation coefficients,
	str(bdn,bdn,bdn)	$\begin{split} \gamma_{(a)(b)(c)} &:= \frac{1}{2} \left(\lambda_{(a)(b)(c)} + \lambda_{(c)(a)(b)} - \lambda_{(b)(c)(a)} \right) \\ &\text{structure constants,} \\ C_{(a)(b)(c)} &:= \gamma_{(a)(c)(b)} - \gamma_{(a)(b)(c)} \end{split}$

TABLE III: Initial set of preprogrammed objects available for basis calculations. All the objects in TABLE I are integrated with basis input and can be run directly from a basis.

		Version 1.0 Initial development by Mustapha Ishak and Kayll Lake
	P	Public Information Help Clear Results Public Input
		Private Database List All Private Input
r		Click to Search Database
		Click to Start Calculator
		Click to Start Calculator Search Results From Database
	Name	
1	Name schwbf	Search Results From Database
		Search Results From Database Description
2	schwbf	Search Results From Database Description 'Null form of Kruskal Szekeres covariant basis'

FIG. 1: GRDB Home Page (at the WWW address http://grdb.org). The operations on the database available at the time this paper was written are displayed, followed by the search form and the Calculator Applet button. Part of the search result by the key word "Kruskal" is displayed.

Record Details for: schwmf

Back to Previous Page

Mark To Load in Calculator

 $ds^2 = -16 \left[(m^2 (2 \text{ m-r(u,v)}) \text{ d u d v }) / \text{uvr(u,v)} \right] + (r(u,v))^2 \text{ d } \theta^2 + (r(u,v))^2 (\sin(\theta))^2 \text{ d } \phi^2$

Metric Name: schwmf

Metric Information: 'Null form of Kruskal Szekeres metric':

Metric References: ["Kruskal, pr, v119, p1743, (1960)", "Szekeres, pmd, v7, p285, (1960)"]

Metric Dimension: 4

Coordinate x1_: u

Coordinate x2_: v

Coordinate x3_: theta

Coordinate x4_: phi

Metric Signature: 2

Metric Complex: ()

Metric Component g12_: -8*m^2*(2*m-r(u,v))/(u*v*r(u,v))

Metric Component g33_: r(u,v)^2

Metric Component g44_: r(u,v)^2*sin(theta)^2

Metric Constraints : [diff(r(u,v),u) = -2*m*(2*m-r(u,v))/(r(u,v)*u), diff(r(u,v),v) = -2*m*(2*m-r(u,v))/(r(u,v)*v)]

Kramer et al. reference: '(13,24) p.158'

Metric Line Element (HTML): -16 $[(m^2(2 \text{ m-r}(u,v)) \text{ d.u.d. v.})/uvr(u,v)] + (r(u,v))^2 \text{ d.}\theta^2 + (r(u,v))^2(\sin(\theta))^2 \text{ d.}\phi^2$

FIG. 2: Details Of The Manifold's Record. The corresponding line element is displayed in a suitable "Tex to Html" mathematical format. The fields contain the components of the metric tensor, the referencing data and an information line on the manifold. The button at the top serves to mark the record to be loaded in the Calculator.

```
Server Load/Select Custom Help Options Coordinate calculations NP Tetrad calculations Basis calculations
Smplifications Translate

GRTensorJ: Version GRDB

Expand and Factor has been calculated.

For the kerrac specetime:

""R_(a b c d ; e) R^(a b c d ; e)": diRiea"

2 4 3 2 2 2

diRiea = -720 M (r + 4 r a cos(theta) - 6 r a cos(theta)

3 3 4 4 3 3

- 4 r a cos(theta) + a cos(theta) ) (r - 4 r a cos(theta)

2 2 2 3 3 4 4 4 3

- 6 r a cos(theta) + 4 r a cos(theta) + a cos(theta) )

2 2 2 2 2 2 2 2 2 2 2 2 3 3 4 4 4 3

- 6 r a cos(theta) + 4 r a cos(theta) + a cos(theta) )
```

FIG. 3: GRTensorJ Graphical User Interface. Calculation of the differential invariant ("diRiem") $R_{abcd;e}R^{abcd;e}$ for the Kerr metric is shown. At the time of writing, this calculation executes in under one second on a contemporary PC. The calculation commands are selected from menus and sub-menus. The help is embedded in the menu and appears under "Explanation".

	grdb Public Line Element Input Page	
Your name:		
e-mail address (required)	
Suggested name	(ø.g. Schwarzschild)	
Reference (jour	al, text, or archive)	
Dimension (e.g	4)	
	g. [x, theta, R, Psi])	
Coordinates (e.		

FIG. 4: GRDB Public Line Element Input Page. The fields are entered by the user following the syntax provided in the examples shown at the top of each box.